

ALMA Observatory Status and new frontiers

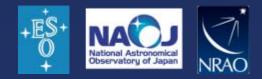
Jorge Ibsen Head of ALMA Department of Computing Joint ALMA Observatory, Santiago de Chile, Chile













Data type

Scientific data

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- - Scientific requests

- 1 Proposal submission
- 2 Observation planing
- 3 Scientific data flow (main archive)

NAC)

ARC EA

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- 4 Scientfic data distribution (ARCs)
- 5 Archive queries
 - ARC: ALMA Regional Center

International partnership

- International partnership among East Asia, Europe and North America in Chile. 20 countries involved
- Total Construction Cost ~USD 1.5 B
- Inaugurated in March 2013

~31 years from since initial idea ~21 years from site hunting

~11 years from starting



Joint ALMA Observatory

- ALAE

CILLES STATES



OSF Hotel Archive Laboratoreo Maintenance Array Control Center Integration Center

AOS Antenas Power Fiber Optic Network Local Oscilator (Timing) Correlator

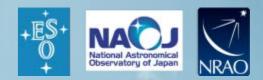
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AOS Technical Building

SCO

Main Archive Data Transmmition to the ALMA Regional Center Offices: -Science -Computing -Administration -Management

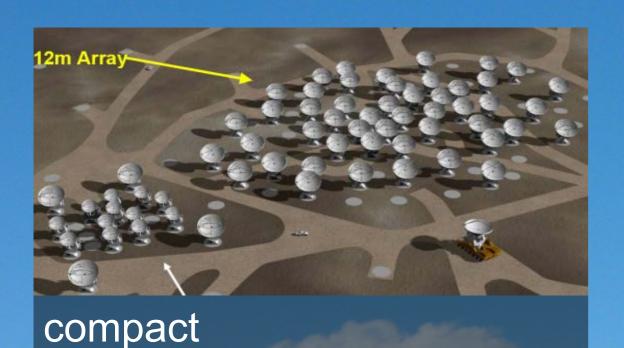




AOS as of May 13, 2014. Photo courtesy of J. Guarda



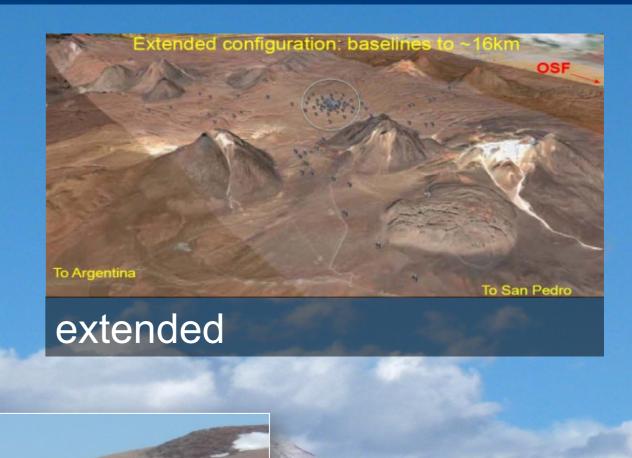
The Atacama Large Millimeter/sub-millimeter Array (ALMA) is array composed of 66 antennas, using *aperture synthesis*, as a "zoom telescope" over the entire accessible mm/submm wavelength range



Built to operate >30 yrs

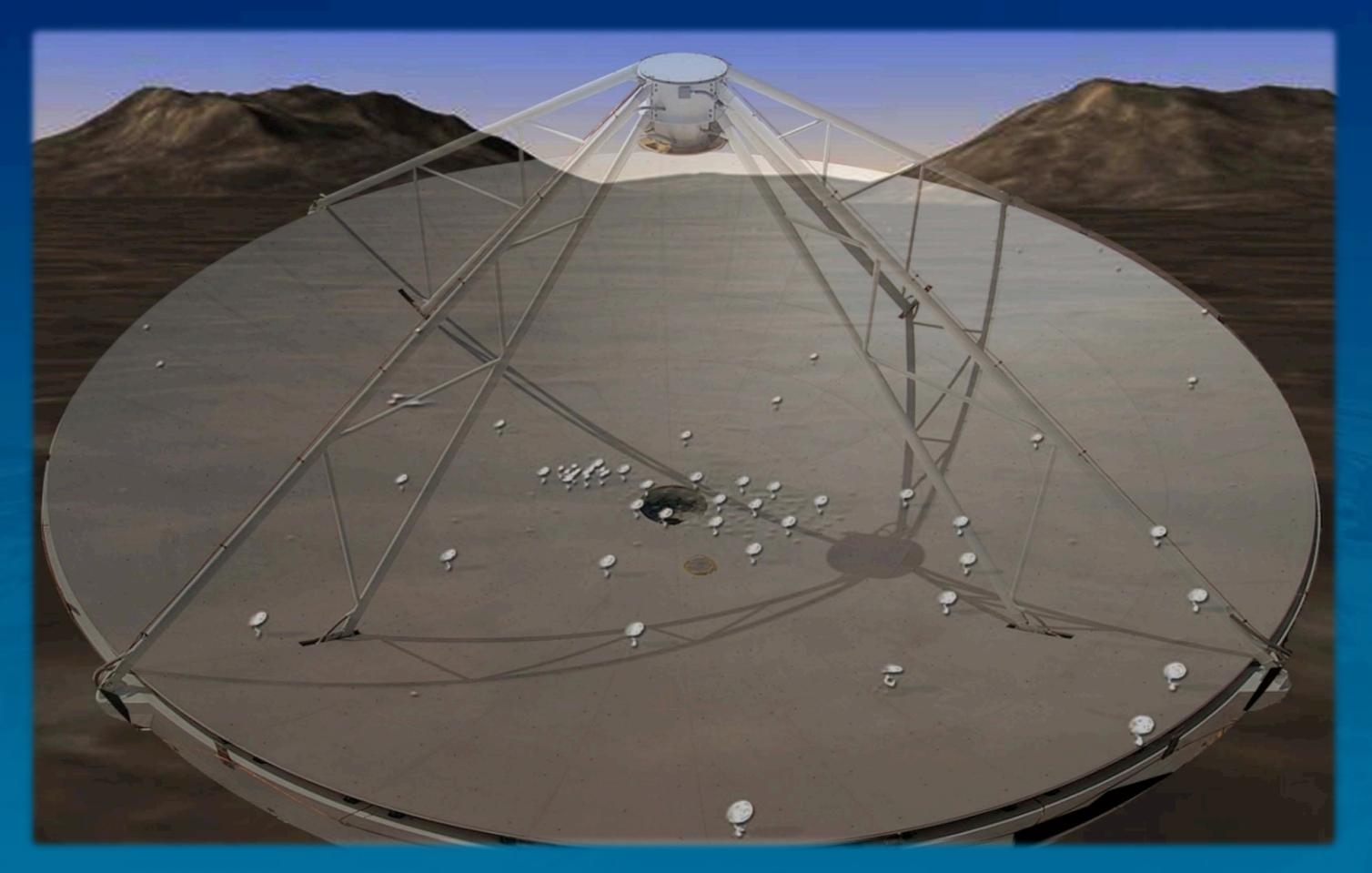
At 5000m





Remotely operated from OSF Control room





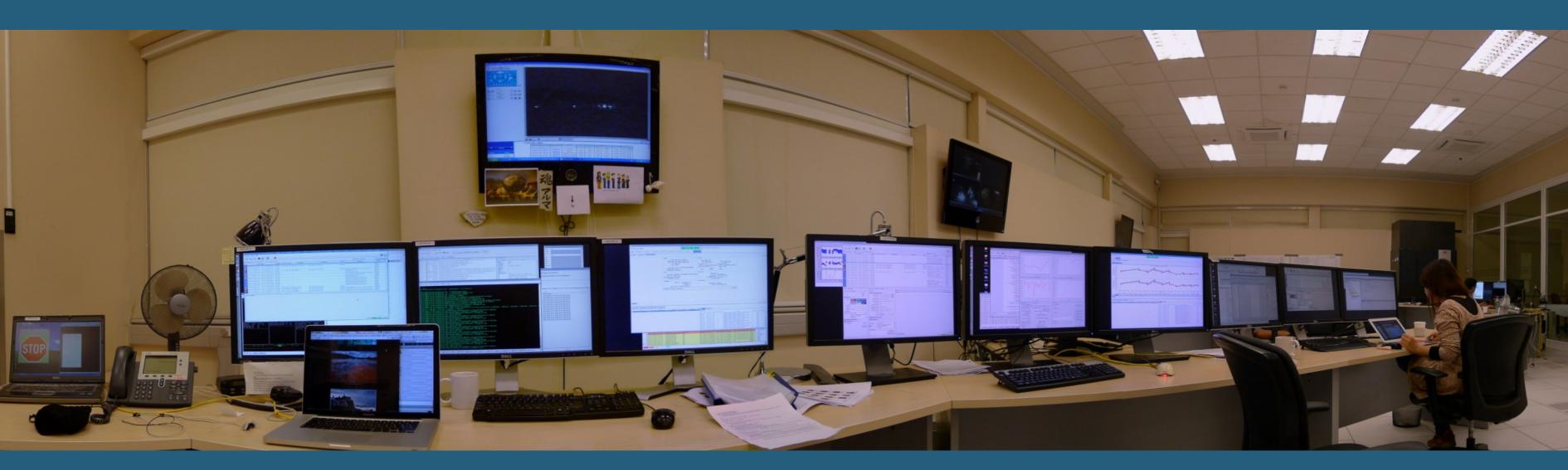
Aperture covered in time by:

- moving the antennas to other foundations (there are 192 of them)
- Earth rotation











Control Room



Data Archive: about 750 GB/day

MIMIM

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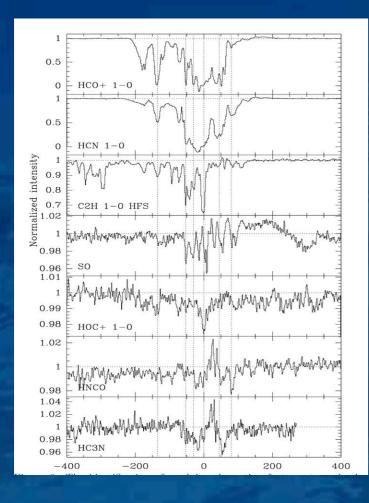


What brings ALMA to Science?



Angular Resolution

♦ better than Hubble ST in the visible



Spectral Resolution

Sensitivity Speed

large increase: 10 -100 times \diamond - 7000m² collecting area - State of the art receivers



\diamond ~10-100 times better than current mm interferometers

$\diamond << 100$ m/s velocity resolution (radio technique!)



Early Science: Cycle 0 March 30, 2011 – January 1, 2013

- \diamond Capabilities: 16 12-m antennas, receiver bands 3, 6, 7 & 9 (wavelengths of about 3, 1.3, 0.8 and 0.45 mm), baselines up to 250m, single field imaging, and a restricted set of spectral modes
- \diamond 919 Proposals: 113 High Priority projects, 51 fillers (500h)





Early Science: Cycle 1 May 31, 2012 – May 31, 2014

Capabilities: 32 12-m antennas in the 12-m Array + 9 7-m antennas, and 2 12-m antennas for total power (single dish) observations, same receiver bands as in cycle 0, baselines ranging from 160 m to 1 km, single field imaging and mosaics of up to 150 pointings, and continuum and spectral line observations

♦ 1131 Proposals: 197 High Priority projects, 92 fillers (800h)
♦ Proposal Review Committee: 80 people met in Santiago





- ♦ Thirty-four 12-m antennas in the 12-m Array, nine 7-m antennas in the 7-m Array, and two 12-m antennas in the TP Array, Receiver bands 3, 4, 6, 7, 8 & 9 (wavelengths of about 3.1, 2.1, 1.3, 0.87, 0.74 and 0.44 mm)
- \diamond Both single field interferometry and mosaics, spectral-line observations with all Arrays (except Band 9 with the TP Array) and continuum observations with the 12-m Array and the 7-m Array, polarization (on-axis, continuum, selected frequencies in Band 3, 6 and 7, no ACA, no mosaics, no circular polarization), mixed correlator modes (both high and low frequency resolution in the same observation)
- \diamond Baselines up to 1 km for Bands 8 & 9, Baselines up to 1.5 km for Bands 3, 4, 6, & 7
 - 1381 Proposals: 353 highest-priority projects

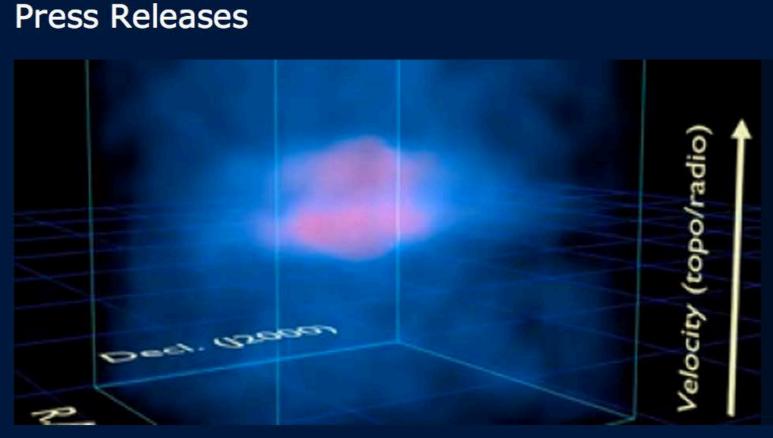




\diamond Transformational science started

 \diamond Increasing number of papers using ALMA data

♦ Frequent announcement of new discoveries available at the ALMA web site





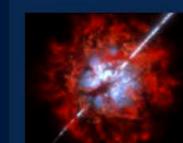
Tuesday, 05 August 2014



Wednesday, 30 July 2014



Thursday, 03 July 2014



Wednesday, 11 June 2014 Bursts

Visit http://www.almaobservatory.org/en/press-room/press-releases

Thursday, 07 August 2014

ALMA Confirms Comets Forge **Organic Molecules in Their Dusty Atmospheres**

ALMA Pinpoints Pluto to Help Guide NASA's New Horizons Spacecraft

ALMA Finds Double Star with Weird and Wild Planet-forming Discs

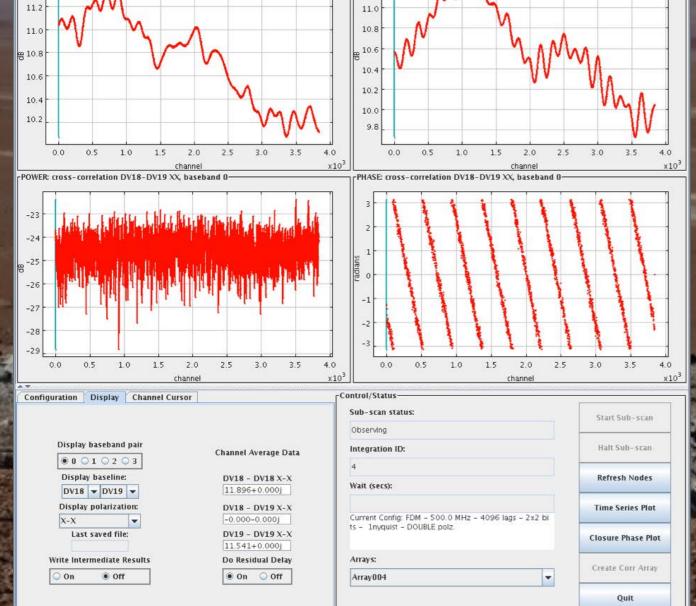
Dynamical Star-Forming Gas Interaction Witnessed by ALMA

Gigantic Explosions Buried in Dust: ALMA Probes Environment around Gamma Ray



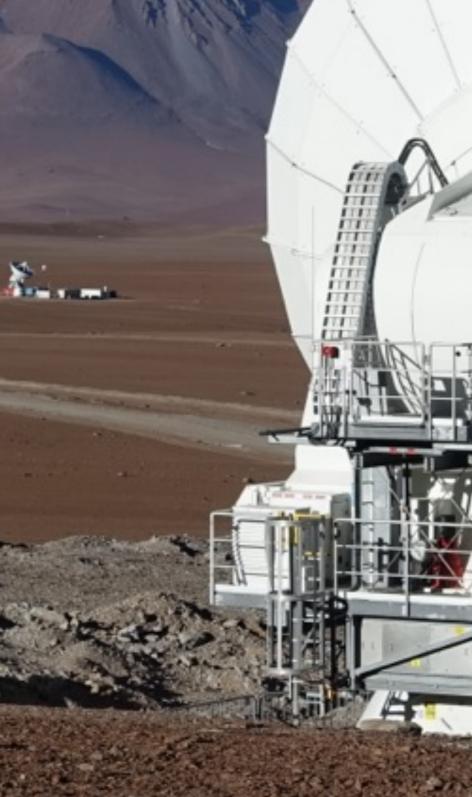
uuto-correlation DV18 XX, baseband (

11.4



11.2

WER: auto-correlation DV19 XX, baseband 0-



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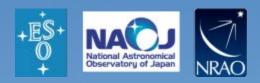
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DV-19

First fringe on a baseline of 7.0425 km obtained on August 24, 2014 at 9:00 pm (CLT) at frequency 92.2 GHz with DV19 on pad P410





ALMA Development Projects for data

ALMA Data Mining Toolkit (ADMIT) ♦ Objective: To increase the scientific productivity of ALMA by creating a set of science-oriented data products and tools which enhance proposers' and archival data users' access to

the science content in ALMA data cubes.

Cube Analysis and Rendering Tool for Astronomy (CARTA) Objective: Provide a next generation drop-in replacement for the current CASA viewer ♦ Both local and remote visualization capabilities ♦ Extensibility and maintainability ♦ Enable Community Development

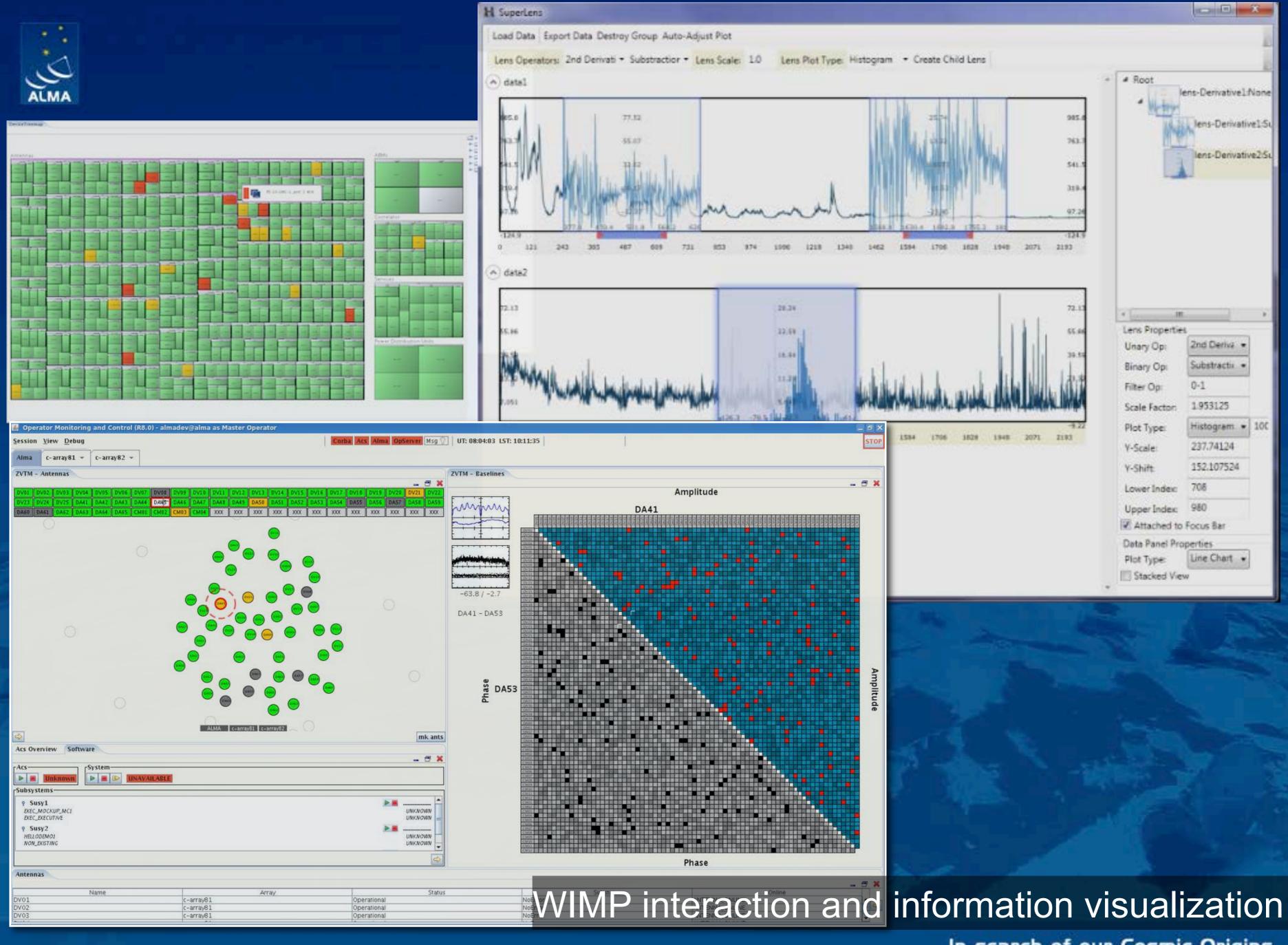




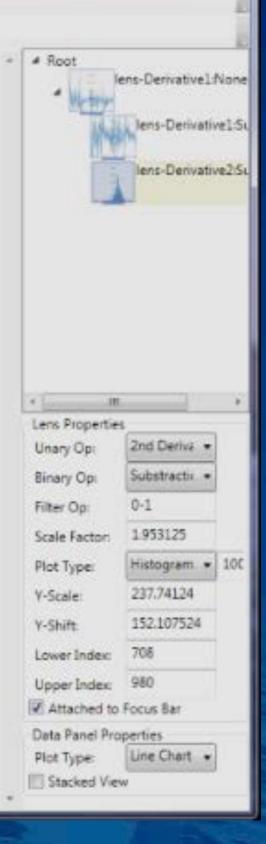
Collaborations: INRIA

Collaboration with Institute National de Recherche en Informatique et en Automatique (INRIA) started informally on May 2009. MoU signed in November 2011 with the objective of *both improving the efficiency and ease of use of ALMA and also of becoming a model state-of-the art development at a world class research facility.*







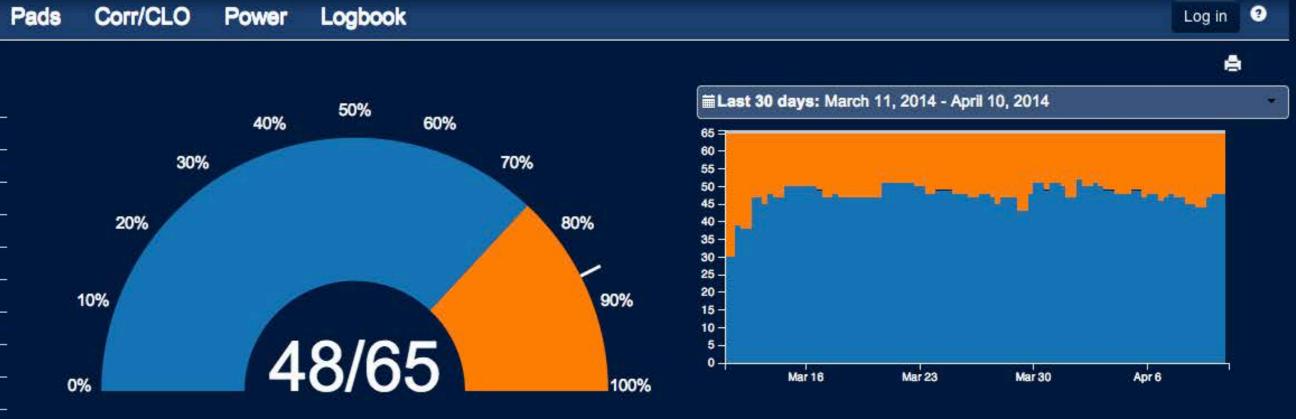


ALMA Dashboard Overview

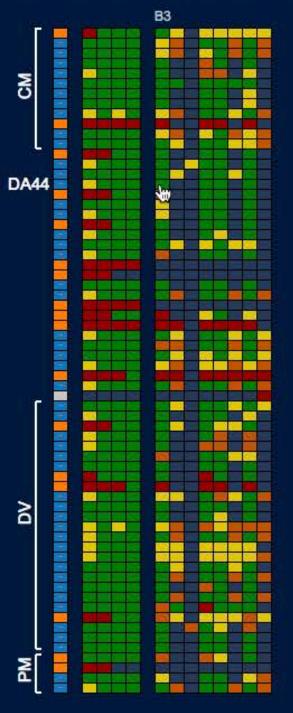
Antennas

✓ Summary

	B 3	B4	B5	B6	B7	B8	B 9	B10
CM Antennas	6	1	0	6	9	1	5	0
Inner Array: Outer Ring	11	0	0	13	13	0	12	0
Southern Arm	0	0	0	0	0	0	0	0
DA+DV Antennas	25	1	0	27	26	1	26	0
Western Arm	0	0	0	0	0	0	0	0
Inner Array: Central Cluster	14	1	0	14	13	1	14	0
Pampa La Bola Arm	0	0	0	0	0	0	0	0
Total	32	2	0	35	37	2	32	0
PM Antennas	1	0	0	2	2	0	1	0
	B3	B4	B5	B6	B7	B8	B9	B10



🗸 Array	Elements	s																	•	
Full	Clear filte	ers															Available Not Availab	e O=Operational C=Science E=Engineering S=	Shutdown	Not Installed
Antenn	a F	Pad	AE	AS	EE	ACD	B3	B4	B5	B6	B7	B8	B9	B10	BL	ACA	 Search ADE	Search CSV	Search Re	marks
	. (AV	AV	AV	AT	AV		AV	AV	AV	AV		AV	AV		AV	AT	A V .	A	v



Current Status Location	Timeline Logbook								
DA44	Array Element -	System -							
	B3 → B4 →	B5 - B6-							
	Correlator Analog Input (CAI)	Baseline: 29							
	ADE (Main technical issues and com	ments):							
	PRTSIR-1894: AZ motor timeout disenge PRTSIR-1676: EL Encoder Head 1 fault (
A068	CSV (Status and Comments):								
	Remarks:								
	Membrane: GORETEX, Solar Filter: N/A								
	Save								

Equip	ment *	A	8D ×
87 -	B8▼	89-	B10 ▼
	ACA: NA		
			Clear
TSIR-46: AOS: DA4	4 mount can not initialize	ed axis (Intermitten	
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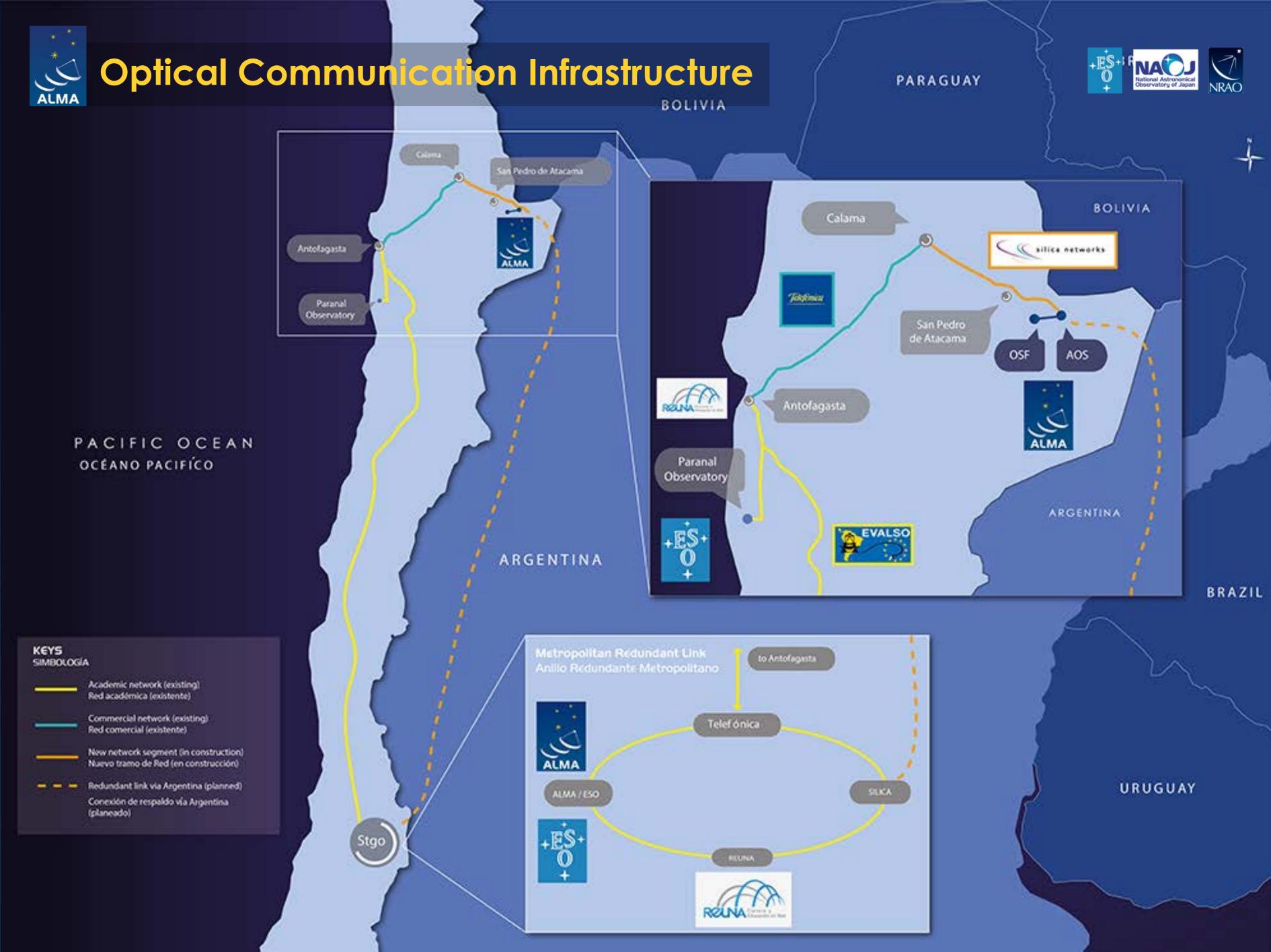


Collaborations: REUNA

MoU signed in January 2012 with the purpose of

- \diamond strongly supporting common efforts to operate and develop high bandwidth connectivity between the different ALMA sites.
- \diamond providing the basis for a joint consideration of the support in connectivity needed by a possible Chilean centre for ALMA data archiving and high performance data reduction.
- ♦ Encouraging informal information exchange and open to future collaborations and projects of mutual interest.







Collaborations: FONDEF Project D11 | 1060

Development of an Astro-Informatic Platform for Management and Intelligent Analisys of Large-scale Data

♦ Main Objective: To develop a virtual observatory, semantic search algorithms on the data (labeled by astronomers), and integration of data mining techniques and artificial intelligence with HPC, allowing large-scale data processing in an efficient and effectively way

♦ Joint project between 5 major universities ♦ Sponsored by FONDEF, REUNA, and ALMA











UNIVERSIDAD DE CHILE











Collaborations: National Laboratory for HPC

NLHPC is a consortia of seven universities and REUNA. Mission

To consolidate a national facility for HPC by offering top quality services and advanced training to answer the national demand for scientific computing, developing links between research groups, the industry and the public sector.

Vision

Participants visualize the NLHPC as a highly competitive center with a range of research services in world-class-quality high performance computing.

MoU signed on January 13, 2014



















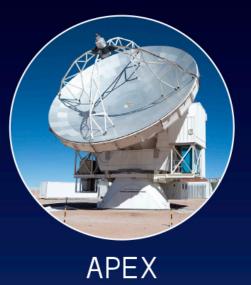




Broader context Observatories in Chile







ALMA



ACT



ASTE



miniTAO



NANTEN2



POLAR BEAR

Broader context Observatories in Chile (Future)



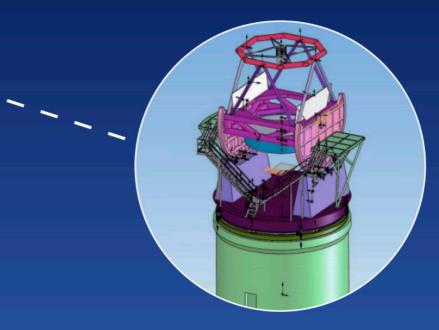
E-ELT



LSST



CCAT



TAO



GMT



Future challenges

- ♦ Produce world-class-top-quality-ready-to-use science data
- Continuous addition of capabilities through development programs
- 24/365 operations with high availability (95% in ALMA's case) while at the same time optimizing (i.e. reducing) operations costs





♦ Infrastructure development \diamond Networks and data centers able to cope with transmission, storage and processing of massive data sets ♦ Research in Astronomy, Computing Science, and Engineering ♦ Archive-based astronomical research \diamond High volume data visualization and data mining ♦ Human Computer Interaction \diamond Network optimization for massive data transmission ♦ Remote operations and virtual presence ♦ Astronomical instrumentation



Thanks!





The Atacama Large Millimeter/submillimeter Array (ALMA), an international astronomy facility, is a partnership of Europe, North America and East Asia in cooperation with the Republic of Chile. ALMA is funded in Europe by the European Organization for Astronomical Research in the Southern Hemisphere (ESO), in North America by the U.S. National Science Foundation (NSF) in cooperation with the National Research Council of Canada (NRC) and the National Science Council of Taiwan (NSC) and in East Asia by the National Institutes of Natural Sciences (NINS) of Japan in cooperation with the Academia Sinica (AS) in Taiwan. ALMA construction and operations are led on behalf of Europe by ESO, on behalf of North America by the National Radio Astronomy Observatory (NRAO), which is managed by Associated Universities, Inc. (AUI) and on behalf of East Asia by the National Astronomical Observatory of Japan (NAOJ). The Joint ALMA Observatory (JAO) provides the unified leadership and management of the construction, commissioning and operation of ALMA.

